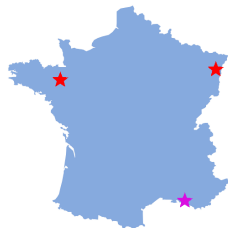


# PICSL : Semi-Lagrangian and Particle Methods for Solving the Vlasov Equation

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- What is plasma ?
- How can we model its dynamics ?
- How can we code a simulation in the chosen model ?
- How can we optimize that code ?

# Examples of plasma

The fourth state of matter... 99% of the universe!

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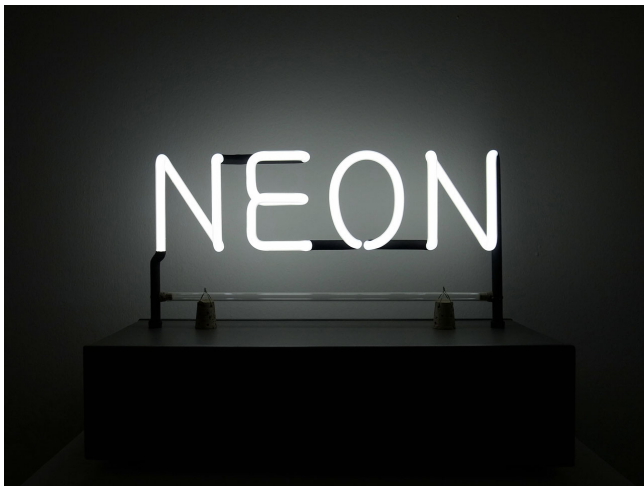
The fourth state of matter... 99% of the universe!



lightning

# Examples of plasma

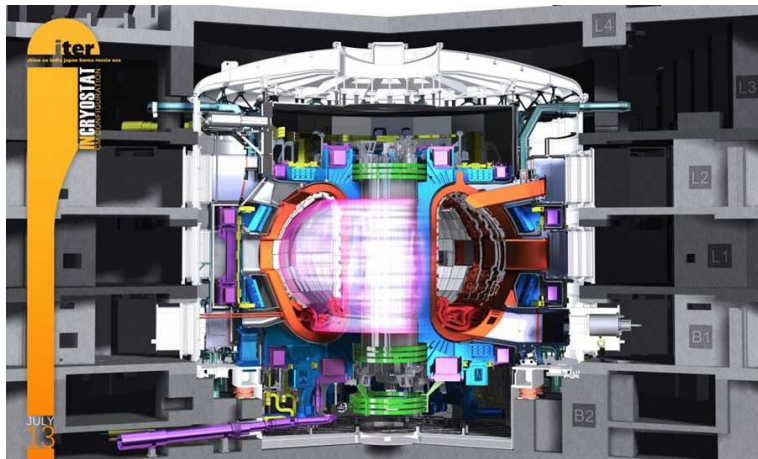
The fourth state of matter... 99% of the universe!



fluorescent light

# Examples of plasma

The fourth state of matter... 99% of the universe !



ITER<sup>a</sup> tokamak (controlled thermonuclear fusion)

a. « The way » (in latin) to produce energy

$$\begin{cases} \frac{\partial f}{\partial t} + \vec{v} \cdot \frac{\partial f}{\partial \vec{x}} - \frac{e}{m} \vec{E} \cdot \frac{\partial f}{\partial \vec{v}} = 0 & \text{Vlasov} \\ -\Delta \phi = \frac{\rho}{\epsilon_0} & \text{Poisson} \end{cases}$$

- $f(\vec{x}, \vec{v}, t)$  : distribution function of the electrons
- $\vec{E}(\vec{x}, t) = -\overrightarrow{\text{grad}} \phi$  : the electric field, here self-induced ;  $\phi$  is the associated scalar potential
- $\epsilon_0$  : vacuum permittivity
- $e, m$  : electron charge and mass
- $t$  : time
- $\vec{x} \in (\mathbb{R}/(L_x\mathbb{Z})) \times (\mathbb{R}/(L_y\mathbb{Z}))$  : particle position (1d, 2d or 3d)
- $\vec{v} \in \mathbb{R}^2$  : particle speed (1d, 2d or 3d)
- $\rho(\vec{x}, t) = e \left( 1 - \int f(\vec{x}, \vec{v}, t) d\vec{v} \right)$  : volume charge density

Essentially three methods for modelling the particle density inside plasma :

- Semi-Lagrangian methods
- Particle-in-Cell methods
- Eulerian methods



- splitting of the Vlasov equation

$$\frac{\partial f}{\partial t} + \vec{v} \cdot \frac{\partial f}{\partial \vec{x}} + \frac{q}{m} \vec{E} \cdot \frac{\partial f}{\partial \vec{v}} = 0$$

- splitting of the Vlasov equation into two simpler equations :

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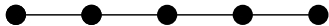
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- follow the characteristics :



→ x

Values after  $k$  time steps.



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Values after  $k + 1$  time steps.

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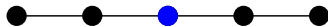
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$g^*(x, (k+1)\delta_t)$

→ x

Values after  $k+1$  time steps.

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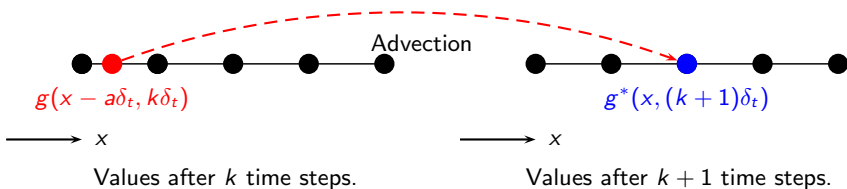
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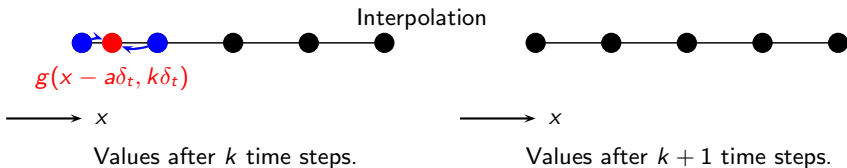
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- approximation of  $f$  via (a lot of) numerical particles
  - one numerical particle represents many real-life particles

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- $$f(\vec{x}, \vec{v}, t) = \sum_{k=1}^N \text{weight}_k \delta(\vec{x} - \vec{x}_k) \delta(\vec{v} - \vec{v}_k)$$

- $\delta$  is the distribution of Dirac :
- $\int_{\mathbb{R}} \delta(x) dx = 1$
- $\delta(0) = +\infty$
- $\delta(x) = 0$  when  $x \neq 0$

- SL

- + only stores  $f$  via a grid on positions and speeds : faster on 1D (2D grid) and 2D (4D grid)
- slower on 3D (6D grid is too much)

- PIC

- + only stores a grid for the fields on positions : faster on 3D
- also stores an array of particles : slower on 1D and 2D
- requires a lot of particles : stochastic convergence in  $\frac{1}{\sqrt{N}}$

Principle : solve exactly the linearized equation.

If  $f^0 \equiv f^0(v)$  is an equilibrium solution and  $f(t=0, x, v) = f^0(v) + A\hat{f}(0, v)e^{ik \cdot x}$  where  $A \ll 1$  then

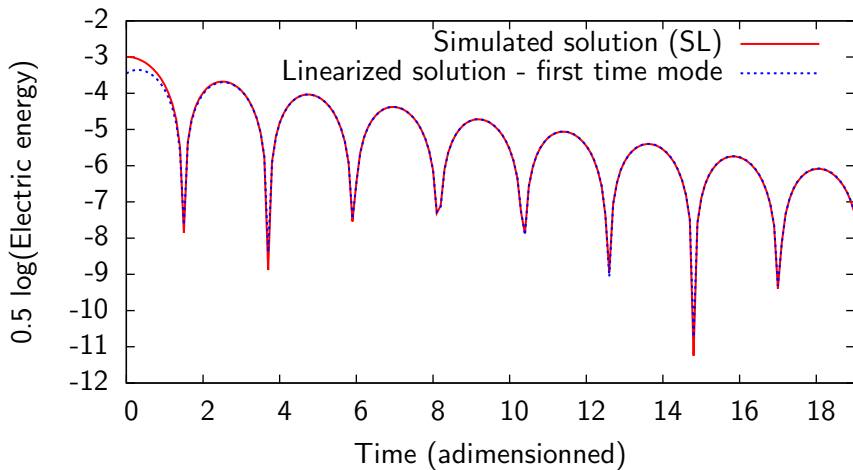
$$E(t, x) = Ae^{ik \cdot x} \sum_{\omega \in D^{-1}(\{0\})} \text{Res}(\omega) e^{-i\omega t} \frac{k}{|k|} + O(A^2),$$

with  $D$  an analytic function depending only on  $f^0$  and  $k$  and  $\text{Res}$  an analytic function depending on  $f^0$ ,  $k$  and  $\hat{f}(0, \cdot)$ .

- $\forall \omega \in D^{-1}(\{0\}), \text{Im}(\omega) \leq 0 \Rightarrow$  stable. e.g.  $f^0 = \frac{e^{-\frac{v^2}{2}}}{\sqrt{2\pi}}$ .
- $\exists \omega \in D^{-1}(\{0\}), \text{Im}(\omega) > 0 \Rightarrow$  unstable. e.g.  $f^0 = \frac{v^2 e^{-\frac{v^2}{2}}}{\sqrt{2\pi}}$ .

# Test case : Landau

$$f(0, x, v) = \frac{e^{-\frac{v^2}{2}}}{\sqrt{2\pi}} (1 + A \cos(kx)) \Rightarrow \forall \omega \in D^{-1}(\{0\}), \mathcal{I}m(\omega) \leq 0$$



Problem : linearly,  $2d$  solution is a superposition of  $1d$  solutions.

Solution : find the term in  $A^2$  in the expansion of  $f$ .

Principle : if  $f = f^0(v) + Af^1(t, x, v) + A^2f^2(t, x, v)$  and  $E = AE^1(t, x) + A^2E^2(t, x)$  then

$$\begin{cases} \partial_t f^2 + v \cdot \nabla_x f^2 - E^2 \cdot \nabla_v f^0 - E^1 \cdot \nabla_v f^1 = 0, \\ -\Delta_x \Phi^2 = -\int_{\mathbb{R}^2} f^2 dv, \\ E^2 = -\nabla_x \Phi^2. \end{cases}$$

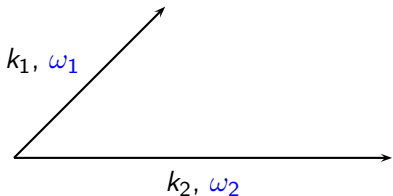
It is the linearized equation but with a source term  $E^1 \cdot \nabla_v f^1$  that is given by the linear analysis. The solution is given by the Duhamel's formula.

Consequence :

- one can deduce the dominant time mode of  $f^2$ .

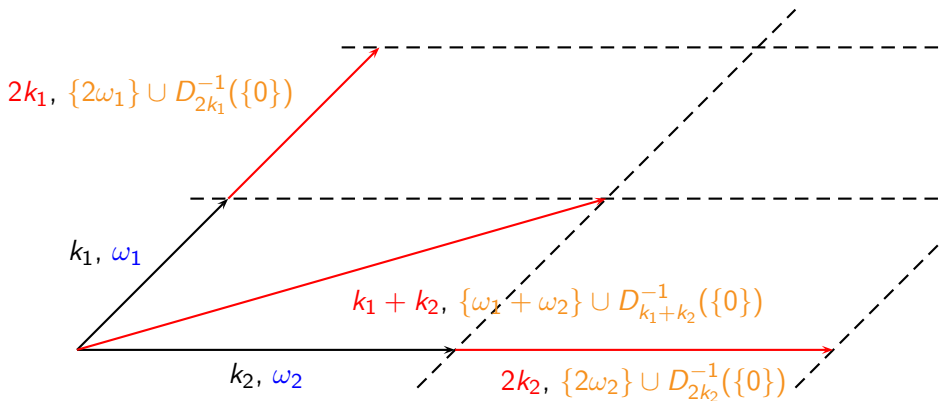
## Test case : $2d \times 2d$

$$f(t = 0, x, v) = f^0(v) + A\alpha(v)e^{ik_1 \cdot x} + A\beta(v)e^{ik_2 \cdot x}$$



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## Test case : $2d \times 2d$

If  $f^0(v) = \frac{v_x^2 e^{-\frac{|v|^2}{2}}}{2\pi}$  and  $L = 4\pi$  then for the spatial modes :

mode  $k$  is unstable  $\Leftrightarrow D_k^{-1}(\{0\}) \cap (\mathbb{R} + i\mathbb{R}_+^*) \neq \emptyset \Leftrightarrow k = \pm \frac{1}{2}(1, 0)$

We take  $f(0, x, v) = (1 + A\cos(\frac{y}{2}) + A\cos(\frac{x+y}{2}))f^0(v)$ ,  $A = 0.001$ .

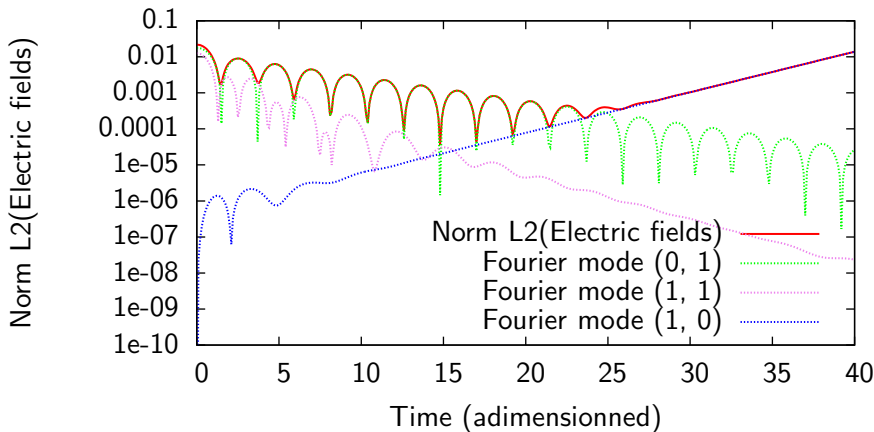
The theory makes us expect :

- a Landau damping at the order 1 in  $A$
- an explosion of the space mode  $\frac{1}{2}(1, 0)$  at the order 2



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# The movie

[http://www.barsamian.am/Slides/2dx2d\\_rho.mpg](http://www.barsamian.am/Slides/2dx2d_rho.mpg)

# Test case : Badsì Herda

Two species,  $\epsilon = \sqrt{\frac{m_i}{m_e}}$ .

Scaling

$$\begin{cases} \partial_t f_i + v \cdot \partial_x f_i + E \cdot \partial_v f_i = 0, \\ \partial_t f_e + \frac{1}{\epsilon} v \cdot \partial_x f_e - \frac{1}{\epsilon} E \cdot \partial_v f_e = 0, \\ \partial_x E = \int_{\mathbb{R}^2} f_i - f_e \, dv. \end{cases}$$

Initial conditions and initialisation :

$$\begin{cases} f_e(t=0, x, v) = \frac{e^{-\frac{v^2}{2}}}{\sqrt{2\pi}}, \\ f_i(t=0, x, v) = 8 \frac{e^{-2v^2}}{\sqrt{2\pi}} (1 + A \cos(kx)). \end{cases}$$

Time modes :

$$D \equiv D_1(\xi) + D_2(\epsilon \xi)$$

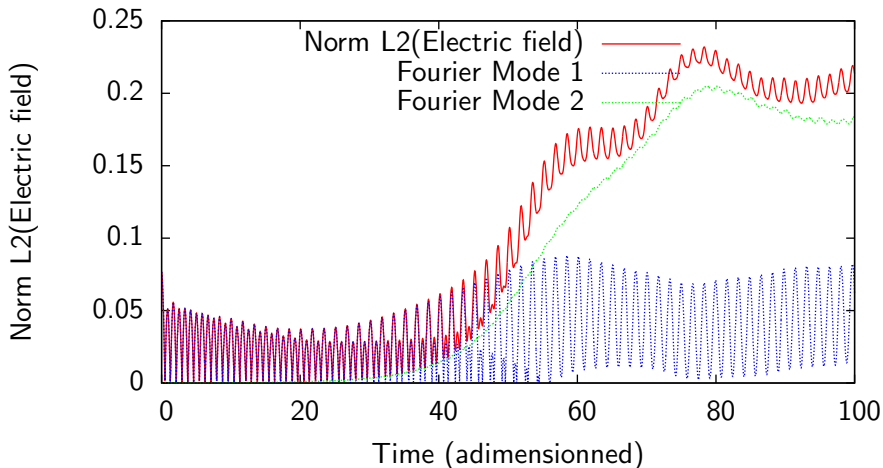
When  $\epsilon \rightarrow 0$ , one has

$$D^{-1}(\{0\}) \equiv [D_1 + D_2(0)]^{-1}(\{0\}) \sqcup \frac{1}{\epsilon} D_2^{-1}(\{0\}) + O(\epsilon).$$

# Test case : Badsi Herda

With  $\omega_1 \in i\mathbb{R}_+^*$  and  $\text{Im}(\omega_2) < 0$  :  $E \equiv Ae^{ikx}(\alpha e^{-i\omega_1 t} + \beta e^{-i\frac{\omega_2}{\epsilon} t})$

Numerically, with  $\epsilon = \sqrt{0.1}$  and  $A = 0.01$ , we have :



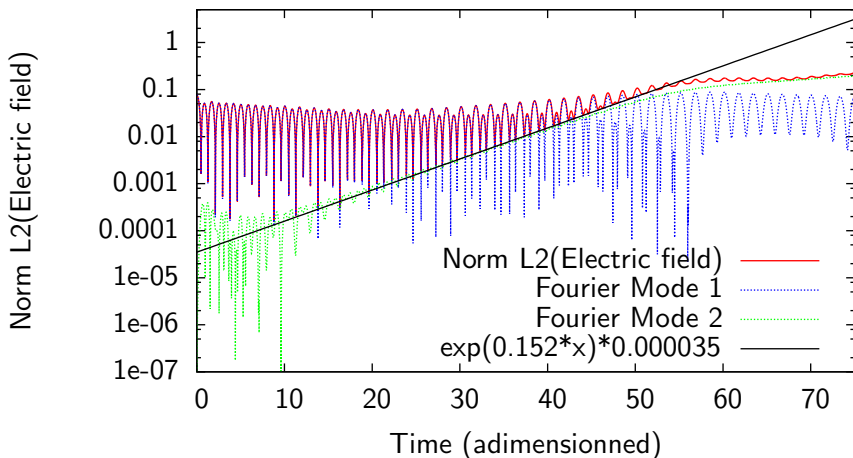
Problem : only the mode  $2k$  explodes.

# Test case : Badsì Herda

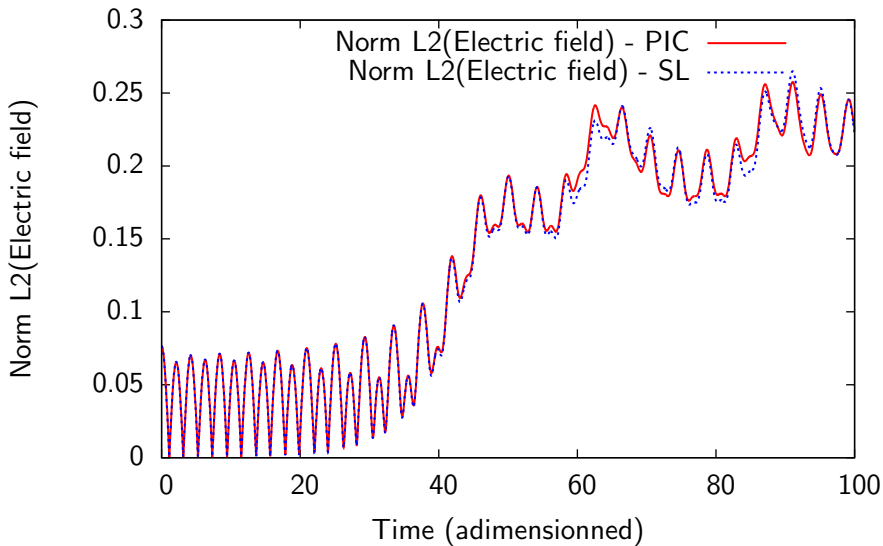
With bilinear analysis, one expects

$$E \equiv Ae^{ikx}(\alpha e^{-i\omega_1 t} + \beta e^{-i\frac{\omega_2}{\epsilon} t}) + A^2 \gamma e^{2ikx} e^{-i2\omega_1 t}.$$

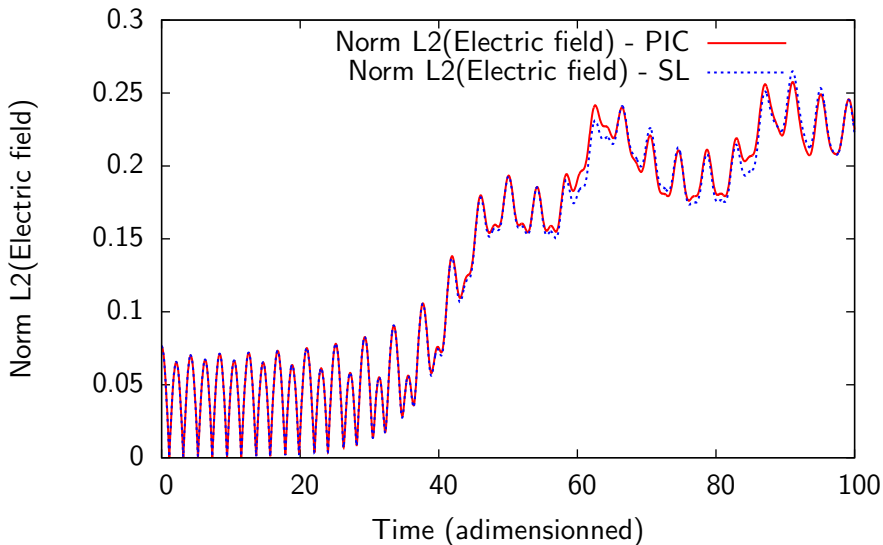
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# Numerical comparison PIC / SL (1D)



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SL : grid  $128 \times 256$ , 2.5h ; PIC :  $8.192 \times 10^9$  particles, 1350h

- extension of the Vlasov-Poisson model to the Vlasov-Maxwell model
- adding of an external magnetic field
- more precise modelling (drift-kinetic)





*That's all Folks!*